

“Experimental Studies on the Chemical Reactivity of Small Oxygenates towards OH radicals at the Ultra-Low temperatures in the Dense Molecular Clouds of the Interstellar Medium”

María Antiñolo,^{1,2} Bernabé Ballesteros,^{1,2} Elena Jiménez,^{1,2} André Canosa,³ Gisèle El Dib³ and José Albaladejo^{1,2}

¹ Universidad de Castilla-La Mancha – Departamento de Química Física – Facultad de Ciencias y Tecnologías Químicas. Edificio Marie Curie. Ciudad Real, 13071 Spain

² Universidad de Castilla-La Mancha – Instituto de Combustión y Contaminación – Camino de Moledores s/n- Edificio Polivalente. Ciudad Real, 13071 Spain

³ Institut de Physique de Rennes. Département de Physique Moléculaire. Campus de Beaulieu, Bât 11C, Université de Rennes I, 35042 Rennes Cedex, France

Abstract

The interpretation of the observed abundances of >180 molecules (ions, radicals and molecules) detected in the interstellar medium (ISM) requires the knowledge of the rate coefficients (k) of all chemical reactions involved in their formation/depletion processes in the gas-phase. Astrochemists usually employ photochemical models where k of a great number of included reactions are not accurately known at the temperatures of dense molecular clouds of the ISM (10-100 K). For many radical-molecule reactions, those models consider the rate coefficient as an estimate or as the extrapolated value from high temperature data at ultra-low temperatures. Extrapolation procedure is not valid for most of the reactions of astrophysical interest, since it has been observed that the reaction rate of radical-molecule reactions is greatly enhanced (even several order of magnitude) at low temperatures. Here, we present the kinetic study of the gas-phase reaction of hydroxyl (OH) radicals with some oxygenated compounds at temperatures of the ISM by using a recently built pulsed CRESU apparatus. The pulsed laser photolysis technique is used to produce OH from H₂O₂ and the laser induced fluorescence technique is employed to monitor the time evolution of OH radicals. The astrophysical implications of the obtained results will be discussed.